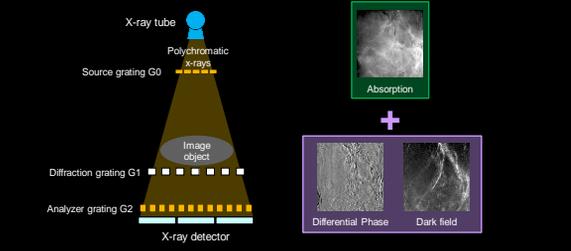


X-Ray Phase Contrast Imaging with Photon-Counting Detectors

Ke Li, PhD
Department of Medical Physics and Department of Radiology
University of Wisconsin-Madison



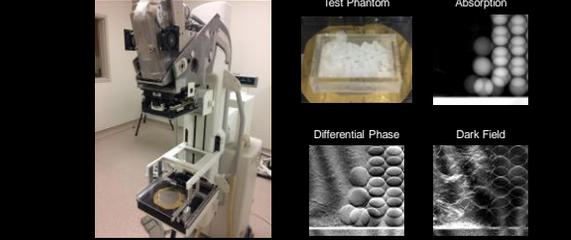
Grating-Based X-Ray Phase Contrast Imaging



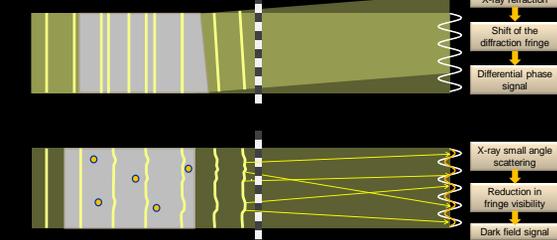
Prototype System at UW-Madison



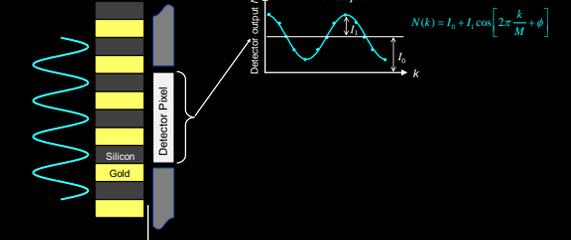
Prototype System at UW-Madison



X-Ray Refraction and Small Angle Scattering



But how does a coarse detector element resolve tiny changes in the diffraction fringe pattern?



But how does a coarse detector element resolve tiny changes in the diffraction fringe pattern?

Without Object: $N(k) = I_0 + I_1 \cos\left[2\pi \frac{k}{M} + \phi\right]$

With Object: $N^{(o)}(k) = I_0^{(o)} + I_1^{(o)} \cos\left[2\pi \frac{k}{M} + \phi + \Delta\phi\right]$

where $\Delta\phi \propto \frac{\partial \phi}{\partial x}$

($\Delta\phi$ is x-ray phase shift induced by the object)

X-Ray Phase Shift

$$\Delta\phi = -\frac{2\pi}{\lambda} \int_L \delta dl$$

$$\delta = \frac{r_0 \lambda^2 \rho_e}{2\pi}$$

$$\Delta\phi = -r_0 \lambda \int_L \rho_e dl = -r_0 \lambda \rho_e L$$

If $L = \frac{\pi}{r_0 \lambda \rho_e}$

$$\Delta\phi = -\pi$$

Legend:
 $\Delta\phi$: x-ray phase shift
 λ : x-ray wave length
 δ : decrement in real part of refractive index
 ρ_e : electron density
 r_0 : classical electron radius
 L : object thickness

π -Phase Diffraction Grating for 32 keV X-Rays

32 keV x-rays

π phase grating designed for 32 keV x-rays

What would happen if the x-ray energy deviate from the designed operating energy?

- The structure height of the grating (L) was set so that for 32 keV x-rays:

$$\Delta\phi_{32\text{keV}} = -r_0 \rho_e L \lambda_{32\text{keV}} = -\pi$$
- Then for the same grating (same L and ρ_e), when the x-ray energy is 16 keV

$$\Delta\phi_{16\text{keV}} = -r_0 \rho_e L \lambda_{16\text{keV}}$$

$$= -r_0 \rho_e L (2\lambda_{32\text{keV}}) = -2\pi$$

π -Phase Diffraction Grating for 32 keV X-Rays 2π -Phase Grating for 16 keV X-Rays

Actual phase shift: -2π

16 keV x-rays

Same grating
Same structure height

Diffraction Pattern with Polychromatic Source

17 keV

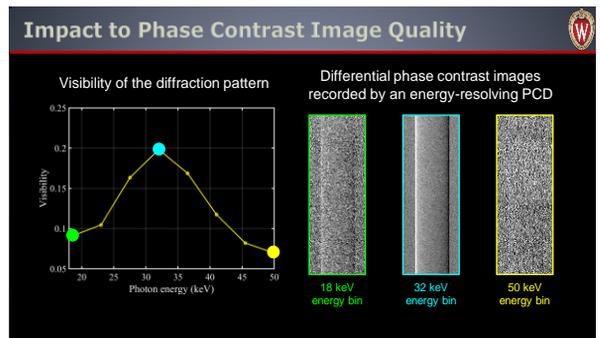
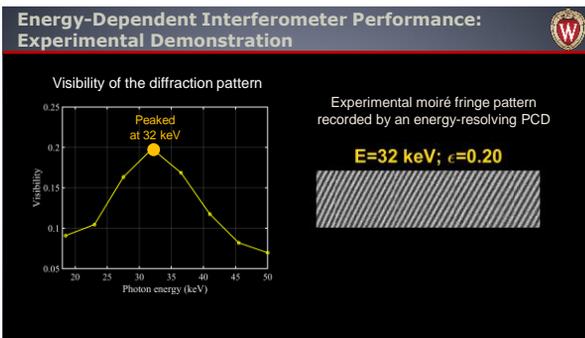
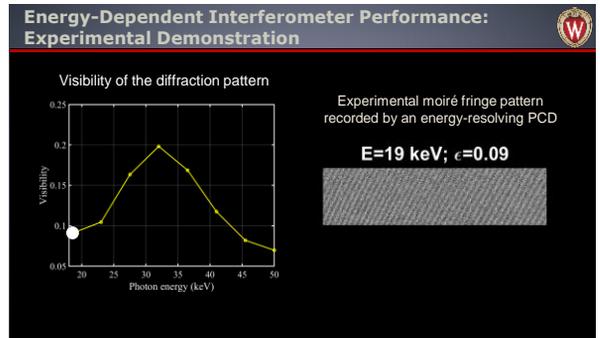
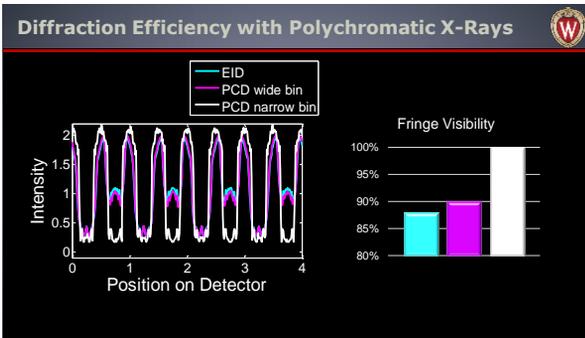
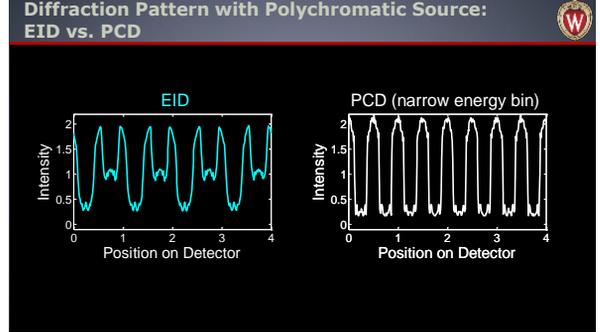
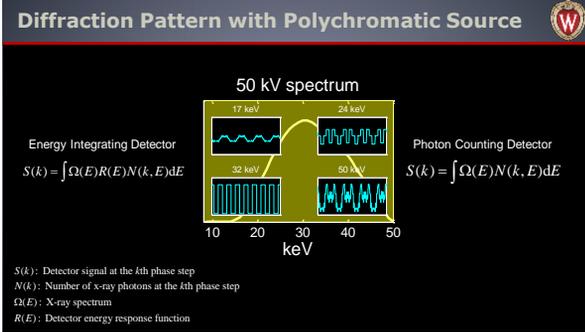
24 keV

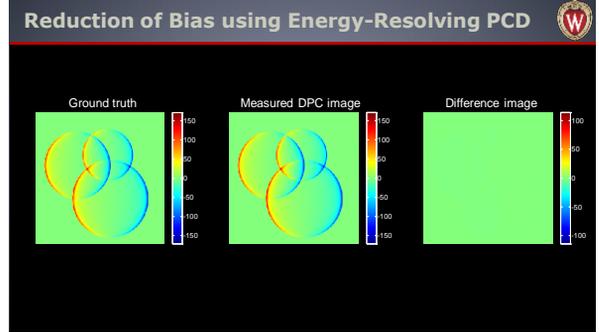
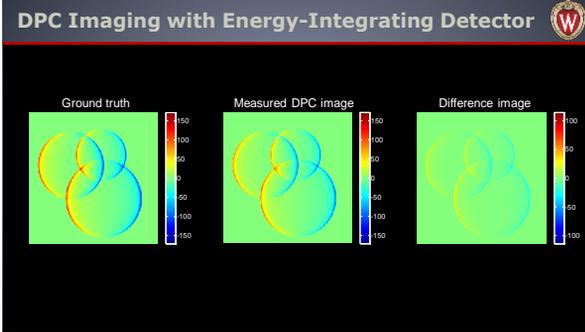
32 keV

50 keV

50 kV spectrum

10 20 30 40 50 keV





Energy Dependence of the Differential Phase Signal

- Differential phase contrast image signal ϕ is related to the energy by

$$\phi(x, y, E) = -\frac{\alpha}{E^2} \frac{\partial}{\partial x} \int \rho_e(x, y, z) dz \quad \dots \dots \dots (1)$$
 where numerical factor α is independent of x-ray energy E .
- Define

$$\varphi(x, y) \equiv -\alpha \frac{\partial}{\partial x} \int \rho_e(x, y, z) dz$$
- Equation (1) becomes

$$\phi(x, y, E) = \frac{1}{E^2} \times \varphi(x, y)$$

Spatial-Spectral DPC Image Matrix

$$\phi(x, y, E) = \frac{1}{E^2} \times \varphi(x, y)$$

Rank (J) = 1

Bin 1	Bin 2	Bin 3
E_1	E_2	E_3

$$J = \begin{bmatrix} \frac{1}{E_1^2} \times \varphi(x_1, y_1) & \frac{1}{E_2^2} \times \varphi(x_1, y_1) & \frac{1}{E_3^2} \times \varphi(x_1, y_1) \\ \frac{1}{E_1^2} \times \varphi(x_1, y_2) & \frac{1}{E_2^2} \times \varphi(x_1, y_2) & \frac{1}{E_3^2} \times \varphi(x_1, y_2) \\ \vdots & \vdots & \vdots \\ \frac{1}{E_1^2} \times \varphi(x_m, y_n) & \frac{1}{E_2^2} \times \varphi(x_m, y_n) & \frac{1}{E_3^2} \times \varphi(x_m, y_n) \end{bmatrix}$$

The number of rows in J is determined by number of image pixels, and the number of columns in J is determined by the number of energy bins

Y. Ge, et al., Vol. 24, Issue 12, pp. 12955-12968, Opt. Exp. (2016)

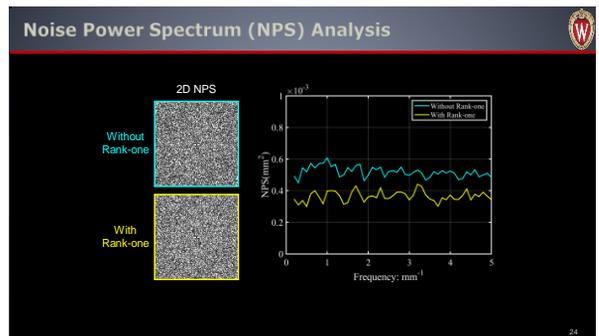
Rank-One Approximation

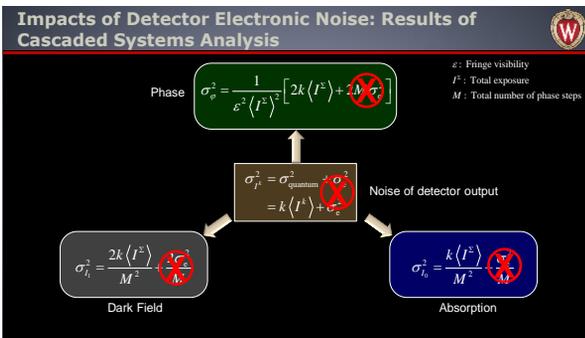
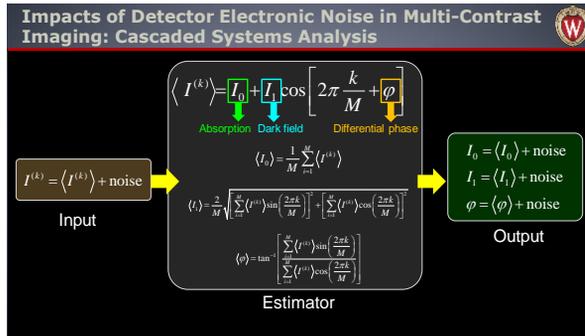
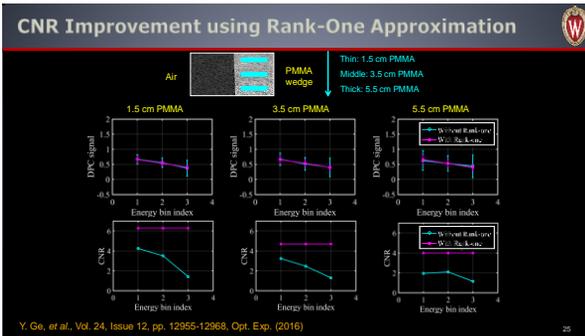
$$J = U \Sigma V^T = \sum_{i=1}^N \sigma_i \mathbf{u}_i \otimes \mathbf{v}_i^T \approx \sigma_1 \mathbf{u}_1 \otimes \mathbf{v}_1^T$$

The wedge phantom provides a constant x-ray refraction angle

	Without Rank-one	With Rank-one
Bin 1		
Bin 2		
Bin 3		

Y. Ge, et al., Vol. 24, Issue 12, pp. 12955-12968, Opt. Exp. (2016)

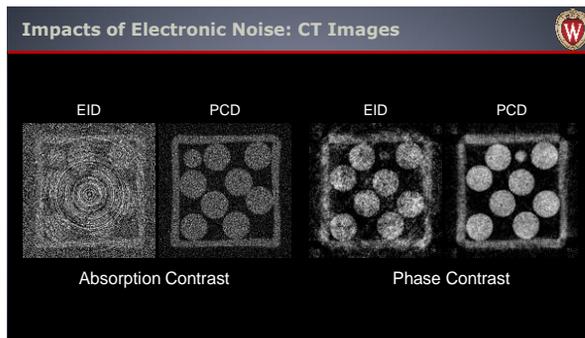
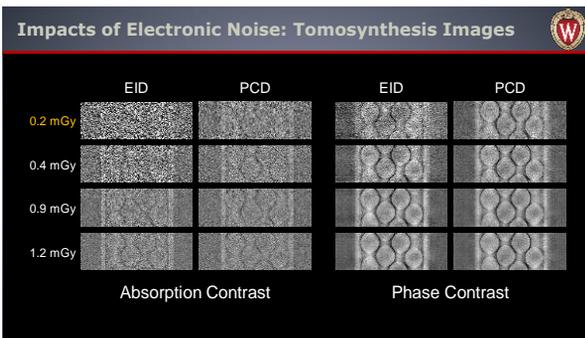




Impacts of Electronic Noise Rejection: Experimental Studies

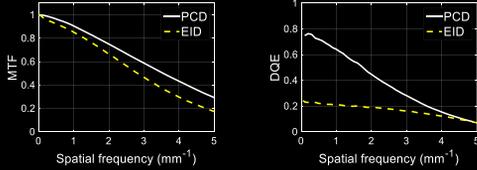
Detector type	Single photon counting	Energy integrating
Model	XCounter XC-FLITE X-1	Shad-o-Box 2048
Pixel size	100 μm	96 μm (after 2 x 2 binning)
Detector active area	15.5 cm x 1.3 cm	10 cm x 5 cm
Conversion method	Direct	Indirect
Conversion material	CdTe	Gd ₂ O ₃ S
Maximum frame rate	1000 fps	2.7 fps
Readout chip	CMOS	CMOS
Bit depth	12	12
XV range	5-300	10-160

X. Ji, et al., "Energy Calibration of Photon Counting Detectors Based On Measurement of X-Ray Attenuation Curve of K-Edge Materials", Wednesday, 7:30 AM - 9:30 AM Room: 601



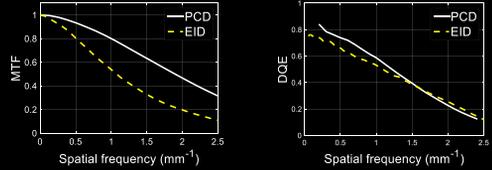
Additional Benefit of PCD: Improvement in Detector DQE(f)

Comparing PCD with EID #1 (thin scintillator):
PCD has slightly better MTF and much better DQE

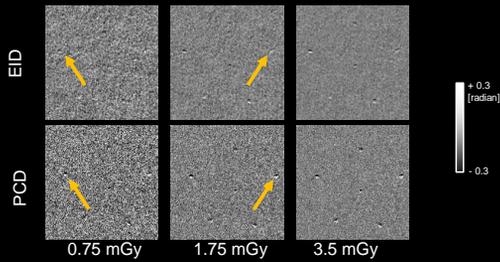


Additional Benefit of PCD: Improvement in Detector Spatial Resolution

Comparing PCD with EID #2 (thick columnar scintillator):
Similar DQE, but PCD has much better MTF

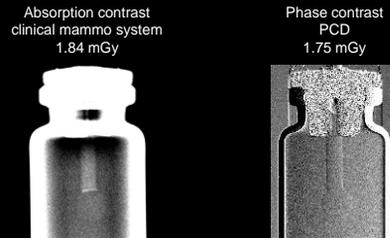


Phase Contrast Images of ACR Mammography Accreditation Phantom



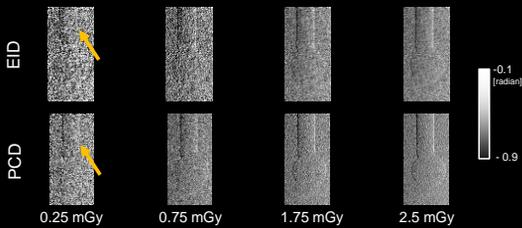
Xu et al., RSNA 2016

Images of In-house Phantom



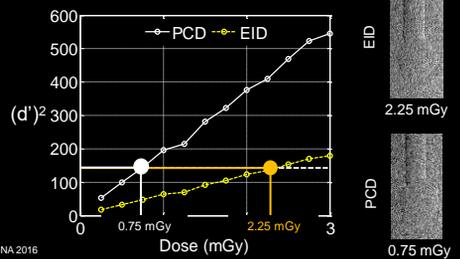
Xu et al., RSNA 2016

Experimental Results of In-house Phantom



Xu et al., RSNA 2016

Improvement in Ideal Observer Performance using PCD



Xu et al., RSNA 2016

Discussion and Conclusion

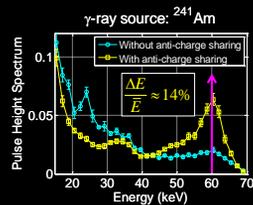
- Performance of x-ray phase contrast imaging (XPCI) is sensitive to x-ray energy
 - Visibility of x-ray diffraction fringes is highest at designed operating energy; it may drop significantly at other energy levels
 - Consequently, the use of polychromatic x-rays decreases fringe visibility
 - Accuracy of the estimated phase contrast signal is highest with monochromatic x-rays
- Performance of XPCI is also sensitive to electronic noise accumulated over multiple phase steps
- Similar to conventional x-ray imaging, performance of XPCI strongly depends on detector DQE

Discussion and Conclusion

- Compared with conventional energy-integrating detectors, photon counting detectors have major advantages in x-ray phase contrast imaging
 - The energy resolving capability of PCD offers the freedom of
 - Selectively utilize a narrow energy window to boost the diffraction efficiency of the grating interferometer, or
 - Jointly utilize all energy windows to boost signal-to-noise ratio
 - Rejection of electronic noise accumulated over multiple phase steps significantly improves phase contrast image quality, especially at low radiation exposure levels
 - Improved DQE offered by PCD leads to better image quality or radiation dose efficiency

Discussion and Conclusion

- The extent to which PCD benefits phase contrast imaging strongly depends on its energy resolution
- Emerging technologies such as anti-charge sharing could provide significant energy resolution improvement
- Further developments in PCD technologies could further improve phase contrast imaging performance



Americium-241; courtesy of UW Cyclotron Group

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Thank You

